Optimization and Control in Shell Models of Turbulence

Forcing optimization of the GOY model

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Content

Context

- Fluid Mechanics
- Navier-Stokes equations
- 2 Main body of work & study
 - Shell Models
 - GOY Model
 - Parameter space & energy
 - Process and complex variable separation
 - Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
 - 4 Conclusions and further work
 - 5 Bibliography

Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Outline

Context

Fluid Mechanics

- Navier-Stokes equations
- 2 Main body of work & study
 - Shell Models
 - GOY Model
 - Parameter space & energy
 - Process and complex variable separation
 - Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Fluid Mechanics

- Conservation of Mass
- Conservation of Momentum
- Energy equation

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Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Turbulence and Laminar Flow

- What is Turbulence and Laminar Flow
- Reynolds number

•
$$Re = \frac{\rho uL}{\mu}$$

Main characteristics



Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Examples of turbulent and fluid dynamics



Figure: Atmosferic flow over Selkirr Island; Laminar to Turbulent flow in cigarette smoke; Jupiter turbulent gases

Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Outline

Context

Fluid Mechanics

Navier-Stokes equations

- 2 Main body of work & study
 - Shell Models
 - GOY Model
 - Parameter space & energy
 - Process and complex variable separation
 - Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Navier-Stokes equations History and context

- Contribuitions made by Claude Navier & George Stokes
- Use and examples of Navier-Stokes equations
- Complexity and possible solutions

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Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

The equations

- Continuity Equation
- Momentum Equations

Three Dimensional components:

$$\begin{split} \nabla \cdot \vec{V} &= 0 \\ \rho \frac{D\vec{V}}{Dt} &= -\nabla p + \rho \vec{g} + \mu \nabla^2 \vec{V} \\ + \text{ boundary conditions } + \text{ inicial conditions} \end{split}$$

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Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Shell Models & Energy

- Structure
- Reasoning behind Shell Models



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Main body of work & study Numerical Integration and Optimization Conclusions and further work Bibliography

Introduction Navier-Stokes equations

Shell Models & NSE

- Kolmogorov 1941 energy spectrum
- Invariants preserved to mimic the Navier-Stokes equations
- Energy cascade spectrum & energy transfer in shell models

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Outline

- 1 Context
 - Fluid Mechanics
 - Navier-Stokes equations

2 Main body of work & study

Shell Models

- GOY Model
- Parameter space & energy
- Process and complex variable separation
- Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Turbulence in Shell Models & NVE

- Fluid Mechanics and Turbulence basics
- Energity Density and Spectral energy Flux
- The four-fifth law
- Parameter Space for the GOY model
- 2D and 3D models

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Kolmogorov Theory

Velocity in the K41

The state of the flow is characterized by the mean energy dissipation per unit mass: $\bar{\varepsilon}$ Velocity difference: $\delta u(l) \equiv |u(r+l) - u(r)|$, eddy of size $l \ll L$; $\delta u(l) = f(\tilde{l}, \bar{\varepsilon})$ by dimensional analysis: $[\delta u] = ms^{-1}$; [l] = m; $[\bar{\varepsilon}] = m^2 s^{-3}$ $[\delta u] = [l]^{\alpha}[\bar{\varepsilon}]^{\beta} \implies \alpha = \beta = 1/3$ Resulting in: $\lambda \delta u(l) = f[\lambda(\bar{\varepsilon}l)^{1/3}]$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Kolmogorov Theory

Scale and Dissipation

With the previous result and the NSE we can estimate the rate of change of the energy per unit volume due to dissipation at scale η , when it becomes important in contrast to L.

$$ar{arepsilon} \sim
u u_i \partial_{jj} u_i \sim
u \delta u(\eta)^2 / \eta^2$$

So the Kolmogorov scale:

 $\eta \sim (\bar{\varepsilon}/\eta^3)^{-1/4}$ depends on the Reynolds numbers as $\eta \sim Re^{-3/4}$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Numeric simulation

The energy cascade can be visualized by setting diferent base parameters.

Figure:



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Spectral energy density

- The mean of the square of the velocity difference is called second order structure function:
- $S_2(I) \equiv \prec \delta u(I)^2 \succ \sim (\bar{\varepsilon}I)^{2/3}$ is related to the spectral energy density through a Fourier transform.
- We can also express the energy density of the flux:

•
$$E = \frac{1}{2} \int u(x)^2 dx = \frac{1}{2} (2\pi)^3 \int_0^\infty u_i(k) u_i(k)^* dk = \frac{1}{2} (2\pi)^3 4\pi \int_0^\infty k^2 |u(k)|^2 dk \equiv \int_0^\infty E(k) dk$$

• Using the Parseval's identity in fourier transforms of spectral velocities:

•
$$\frac{1}{(2\pi)^3} \int u_i(x) u_i(x) dx = \int u_i(-k') u_i(k') dk' = \int u_i(k') u_i(k')^* dk'$$
 with $u_i(x) \in \mathbb{R} \implies u_i(-k') = u_i(k')^*$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Spectral energy density

- Integration was performed over the sphere and absorving the spacial variable into dx, results in:
 - $E(k) = (2\pi)^4 k^2 | u(k) |^2$ and $S_2(l) = \int [u(l+x) u(x)]^2 dx$
 - Combining both e quations and after several complex integration processes we obtain:
 - $E(k) = \frac{1}{2\pi k} \int_0^\infty x \sin(x) S_2(x/k) dx \sim \bar{\varepsilon}^{-2/3} k^{-5/3}$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Spectral energy density



Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

The four-fifth law

• A special case occurs for a third order struture function, associated with energy conservation. Again considering a homogeneus and isotropic turbulent fluid left alone after intense stirring. As the kinetic change is 0, the mean energy dissipation is:

• $\bar{\varepsilon} = -\frac{1}{2}\partial_t \left[u \right]$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Scalling relation

After several integrations and calculation regarding isotropy and the NSE, remains a simple equation regarding the third order structure function:

 $S_3 = -\frac{4}{5}\bar{\varepsilon}I$

This law is a cornerstone of fluid dynamics. Not only is it in agreement with the phenomenology described in the K41 as it is exact and among few exact results derived from NSE.

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Outline

Context

- Fluid Mechanics
- Navier-Stokes equations

2 Main body of work & study

- Shell Models
- GOY Model
- Parameter space & energy
- Process and complex variable separation
- Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Gledzer & Obukhov

Model and given inicial conditions

•
$$\mathcal{L}u = f_n$$

•
$$\mathcal{L}u = \frac{du_n}{dt} - ik_n \left(u_{n+1}u_{n+2} - \frac{\varepsilon}{q}u_{n+1}u_{n-1} + \frac{\varepsilon-1}{q}u_{n-2}u_{n-1}\right)^* + \nu k_n^2 u_n$$

- Constants used
- Force and solution for the GOY model

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Outline

Context

- Fluid Mechanics
- Navier-Stokes equations

2 Main body of work & study

- Shell Models
- GOY Model

• Parameter space & energy

- Process and complex variable separation
- Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Parameter Space for the GOY model

- Number of Shells
- Kinematic viscosity
- Initial force
- Free Parameters

Values used

$$N = 24; \nu = 10^{-7}; f_n = (1 + i) \times 5 \times 10^{-3}$$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

2D and 3D models

- $\bullet~2D/3D$ selector
- Complex non linear coeficients

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Outline

Context

- Fluid Mechanics
- Navier-Stokes equations

2 Main body of work & study

- Shell Models
- GOY Model
- Parameter space & energy

• Process and complex variable separation

- Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Constant manipulation and restraints

Constants changed

GOY model constant substituiton

•
$$a_n = k_n = k_0 q'$$

•
$$b_n = -\frac{1}{2}k_{n-1}$$

•
$$c_n = -\frac{1}{2}k_{n-2}$$

- All components with shell index, $n \leqslant 0$ are taken as zero
- Resulting in:

$$\underbrace{\left(\frac{d}{dt} + vk_n^2\right)u_n = i\left(a_nu_{n+1}u_{n+2} + b_nu_{n+1}u_{n-1} + c_nu_{n-1}u_{n-2}\right)^* + f_n}_{4}$$

Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Separation

• Splitting complex variables $u_n = x_n + iy_n$:

Real Part

$$\left(\frac{d}{dt} + vk_n^2\right)x_n = a_n\left(x_{n+1}y_{n+2} + x_{n+2}y_{n+1}\right) + b_n\left(x_{n-1}y_{n+1} + x_{n+1}y_{n-1}\right) + c_n\left(x_{n-2}y_{n-1} + x_{n-1}y_{n-2}\right) + f_x$$

Imaginary Part

$$\left(\frac{d}{dt} + vk_n^2\right)y_n = ia_n\left(x_{n+1}x_{n+2} - y_{n+2}y_{n+1}\right) + ib_n\left(x_{n-1}x_{n+1} - y_{n+1}y_{n-1}\right) + ic_n\left(x_{n-2}x_{n-1} - y_{n-1}y_{n-2}\right) + if_y$$

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Verification

• Jupyter Notebook, Python, algorithm used to verify the separated variables equations.

Algorithm

- Load a complex variable with pseudo-random numbers (u = x + iy);
- Write the complex GOY model in simple for loops, all the N=24 shells;
- Write the separated GOY equations as for loops;
- Write a function that retrieves the diference between the 3 functions;
- Return the biggest difference(0 meaning these equations were well determined).

Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Outline

Context

- Fluid Mechanics
- Navier-Stokes equations

2 Main body of work & study

- Shell Models
- GOY Model
- Parameter space & energy
- Process and complex variable separation
- Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
- 4 Conclusions and further work
- 5 Bibliography

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 Context
 Shell Models

 Main body of work & study
 GOY Model

 Numerical Integration and Optimization
 Parameter space & energy

 Conclusions and further work
 Process and complex variable separation

 Bibliography
 Linearization of the equations



Linearized the system of real and imaginary equations around set solutions.

Simple example

$$\begin{bmatrix} x_n \\ y_n \end{bmatrix} = \begin{bmatrix} \tilde{x}_n \\ \tilde{y}_n \end{bmatrix} + \varepsilon \begin{bmatrix} x_n \\ y_n \end{bmatrix}$$
The term of order ε^2 will not be considered ($\varepsilon \ll 1$)

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Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Linear Scheme



where the block matrices C_1 , C_2 , C_3 , C_4 , are given by:

Shell Models GOY Model Parameter space & energy Process and complex variable separation Linearization of the equations

Coeficient Part Matrix



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Understanding the Optimization Problem Runge-Kutta methods

Outline

Context

- Fluid Mechanics
- Navier-Stokes equations
- 2 Main body of work & study
 - Shell Models
 - GOY Model
 - Parameter space & energy
 - Process and complex variable separation
 - Linearization of the equations
- 3 Numerical Integration and Optimization
 - Understanding the Optimization Problem
 - Runge-Kutta methods
 - 4 Conclusions and further work
- 5 Bibliography

Understanding the Optimization Problem Runge-Kutta methods

Adjoint operator

- Cost functional in the separated variables equations
- Determining the adjoint operator and transpose matrix

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Understanding the Optimization Problem Runge-Kutta methods

Outline

1 Context

- Fluid Mechanics
- Navier-Stokes equations
- 2 Main body of work & study
 - Shell Models
 - GOY Model
 - Parameter space & energy
 - Process and complex variable separation
 - Linearization of the equations
- Output: State of the state o
 - Understanding the Optimization Problem
 - Runge-Kutta methods
 - 4 Conclusions and further work
- 5 Bibliography

Understanding the Optimization Problem Runge-Kutta methods

Classic fourth order RK

Generic expression

$$u^{n+1} = u^n + \Delta t \sum b_i k_i$$

Inicial Value Problem

$$y' = f(t, y); y(t_0) = y_0$$

$$y_{n+1} = y_n + \frac{h}{6}(k_1 + 2k_2 + 2k_3 + k_4)$$

$$t_{n+1} = t_n + h$$

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Understanding the Optimization Problem Runge-Kutta methods

Classic fourth order RK & GOY model

RK4

$$k_{1} = f(t_{n}, y_{n})$$

$$k_{2} = f(t_{n} + \frac{h}{2}, y_{n} + \frac{h}{2}k_{1})$$

$$k_{3} = f(t_{n} + \frac{h}{2}, y_{n} + \frac{h}{2}k_{2})$$

$$k_{4} = f(t_{n} + h, y_{n} + hk_{3})$$

$$h = 10^{-4}; n \approx 2.5 \times 10^{8}$$

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Understanding the Optimization Problem Runge-Kutta methods

Computing the problem and force optimization

Optimization Process

- Choosing an initial guess $f^{(0)}$; n = 0
- Solve the GOY model equations, with $f = f^{(0)}$
- Solve the adjoint operator from matrix C1...
- Obtain the cost functional gradient
- Find a parameter...
- Update the variable and repeat.

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Understanding the Optimization Problem Runge-Kutta methods

Finding the particular solution

Process

$$\left(\frac{d}{dt} + vk_n^2\right)u_n = i\left(a_nu_{n+1}u_{n+2} + b_nu_{n+1}u_{n-1} + c_nu_{n-1}u_{n-2}\right)^* + f_n$$

- A solution is calculated for every time step.
- A solution consists of 24 pairs of real and imaginary numbers.

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Understanding the Optimization Problem Runge-Kutta methods

Finding other solutions and changing the inicial force.

Plugging the solutions $x_n y_n$ and using the Runge Kutta method for integration, we find solutions for the values $\widetilde{x_n} \& \widetilde{y_n}$. Python algorithm was used to find these solutions.

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Algorithm

Using the GOY separated equations x_n ; y_n are found for a certain δt . Using RK4 Using the linearized equations, and matrices C1, C2, C3, C4 for each time step: $\widetilde{x_n}$ and $\widetilde{y_n}$ are calculated Repeat the process Expected iterations around 2×10^8

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Conclusions

The force optimization process and the verification of convergence for the solutions was not fully completed, mostly due to lack of time, but also to some limitations in Python. It would have been better to do the long computing algorithms in a simpler language such as Fortran. The objective of the long numeric runs was to find the optimal force or aproaching it, optimal in a way that would recreate and verify the energy cascade found in the GOY model and in the K41.

Bibliography

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